

J1105+4444

11:05:08.16
+44:44:47.4

z_{spec}	0.02154 [0.00001]
Hubble distance (CMB)(Mpc)	98.64 ± 6.91
Diameter (")	40.33
Diameter (kpc)	19.79
Object type	Galaxy
Morphology	Irr

SIMBAD : <https://simbad.cds.unistra.fr/simbad/sim-id?ident=%404699439&Name=%5bBKD2008%5d%20WR%20276&submit=submit>
 NED : https://ned.ipac.caltech.edu/byname?objname=MRK+0162&hconst=67.8&omegam=0.308&omegav=0.692&wmap=4&corr_z=1

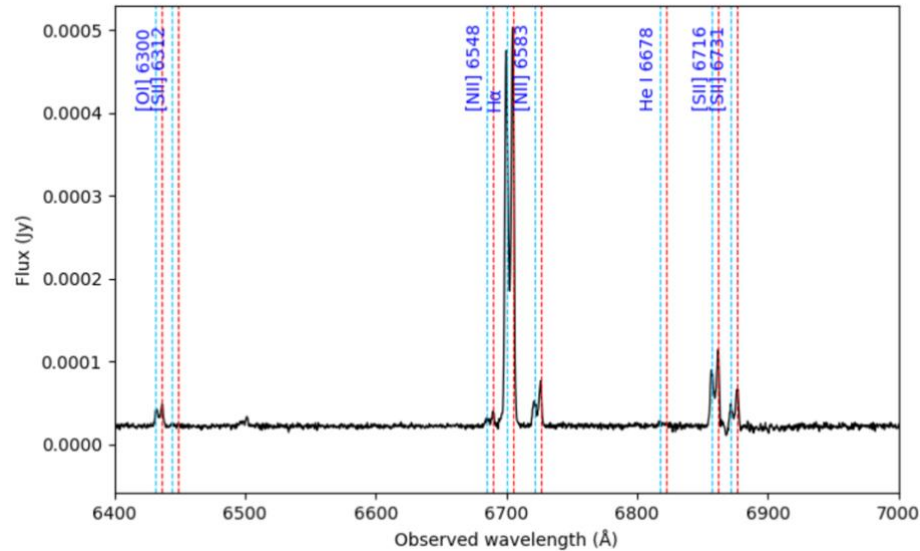
12+log(O/H)	Log M	Log SFR	Log sSFR	O32	$n_e[\text{SII}]$	$T_e[\text{OIII}]$	E(B-V)
8.12	8.69	0.56	-8.13	2.66	141	10680	0.16

Derived properties from CLASSY (<https://www.danielleaberg.com/classy>, target 35 of the sample)

This galaxy exhibits two distinct regions from a spectroscopic perspective: one region where the spectra display single-peaked lines (A System) and another region where the spectra feature double-peaked lines. The new object responsible for the second peak has been named the B System. The calculated redshifts for each are as follows:

$$z_A = 0.02166$$

$$z_B = 0.02088$$

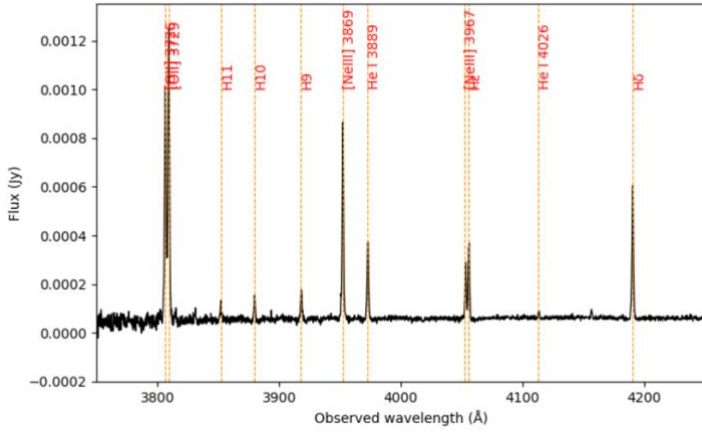


Spectrum with double-peaked lines in LR-R (spaxel 429)

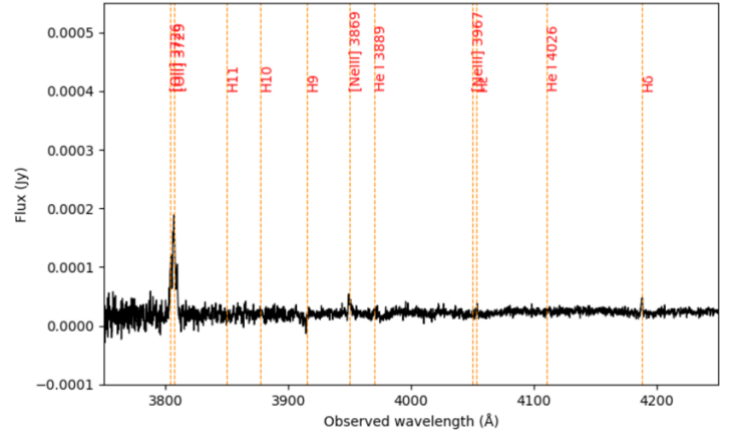
Spectra from the brightest spaxels

LR-U

A system (402)

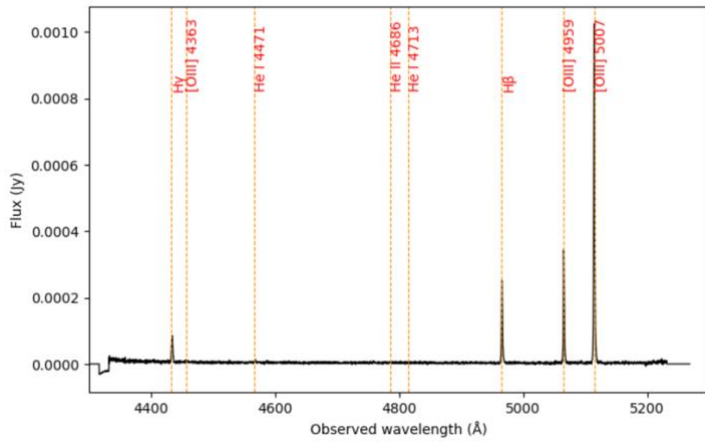


B system (428)

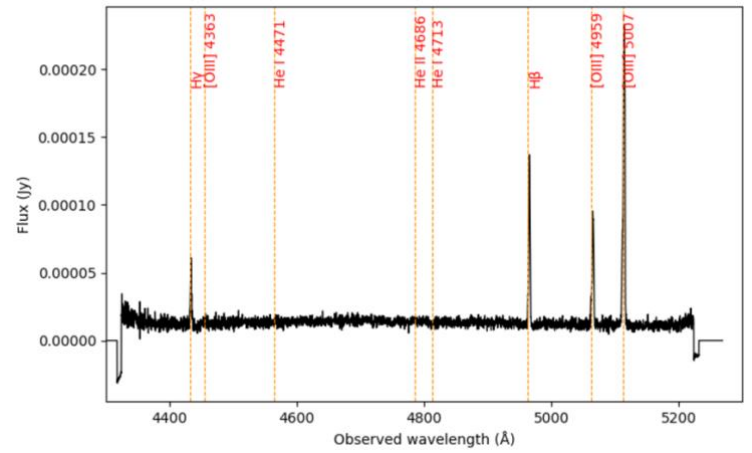


LR-B

A system (291)

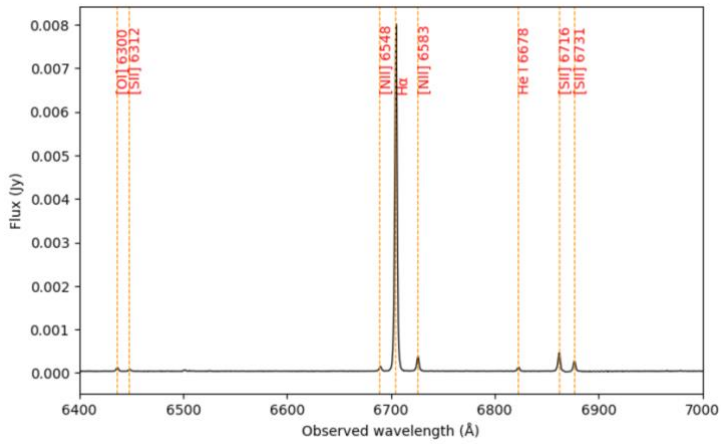


B system (141)

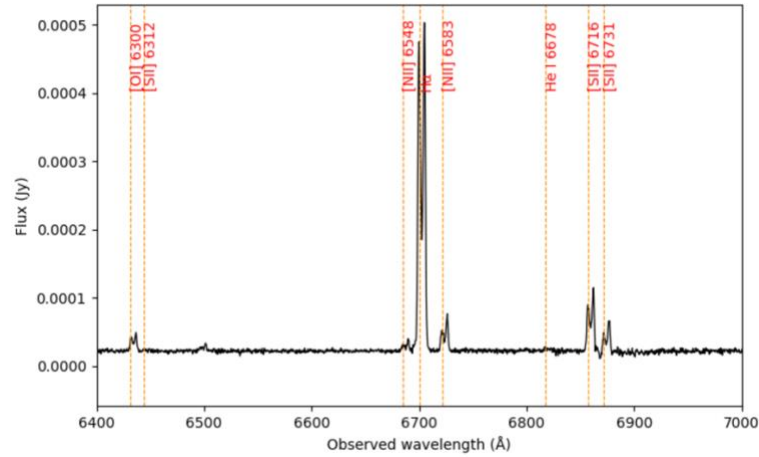


LR-R

A system (289)



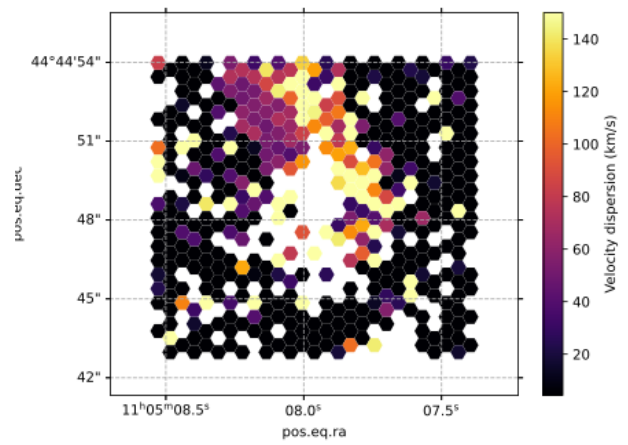
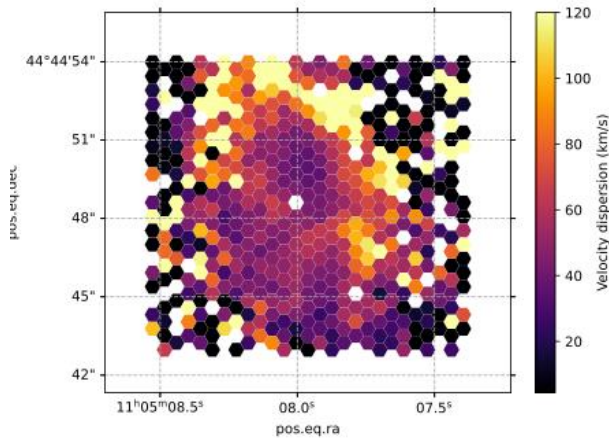
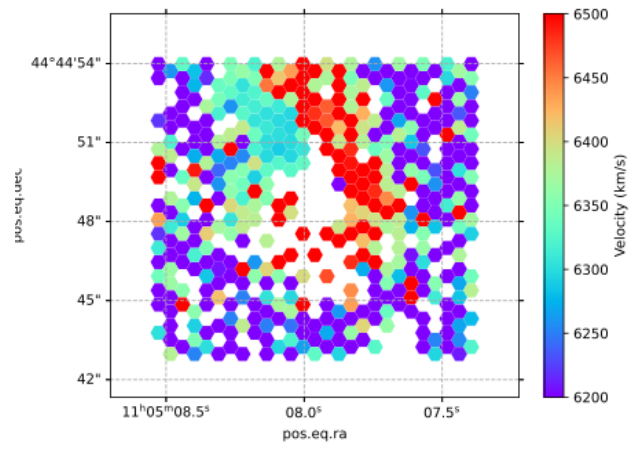
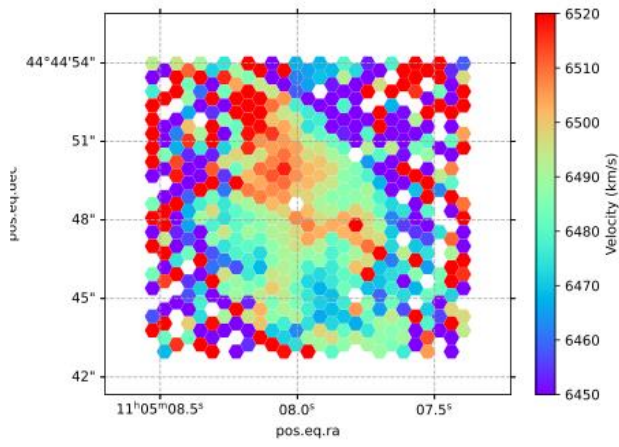
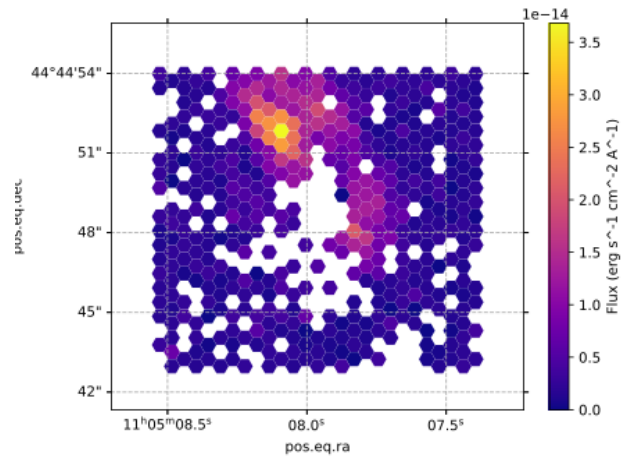
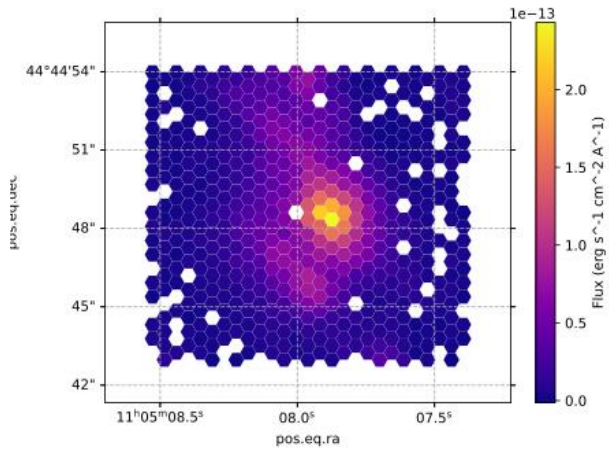
B system (532)



[OII]3726

A system

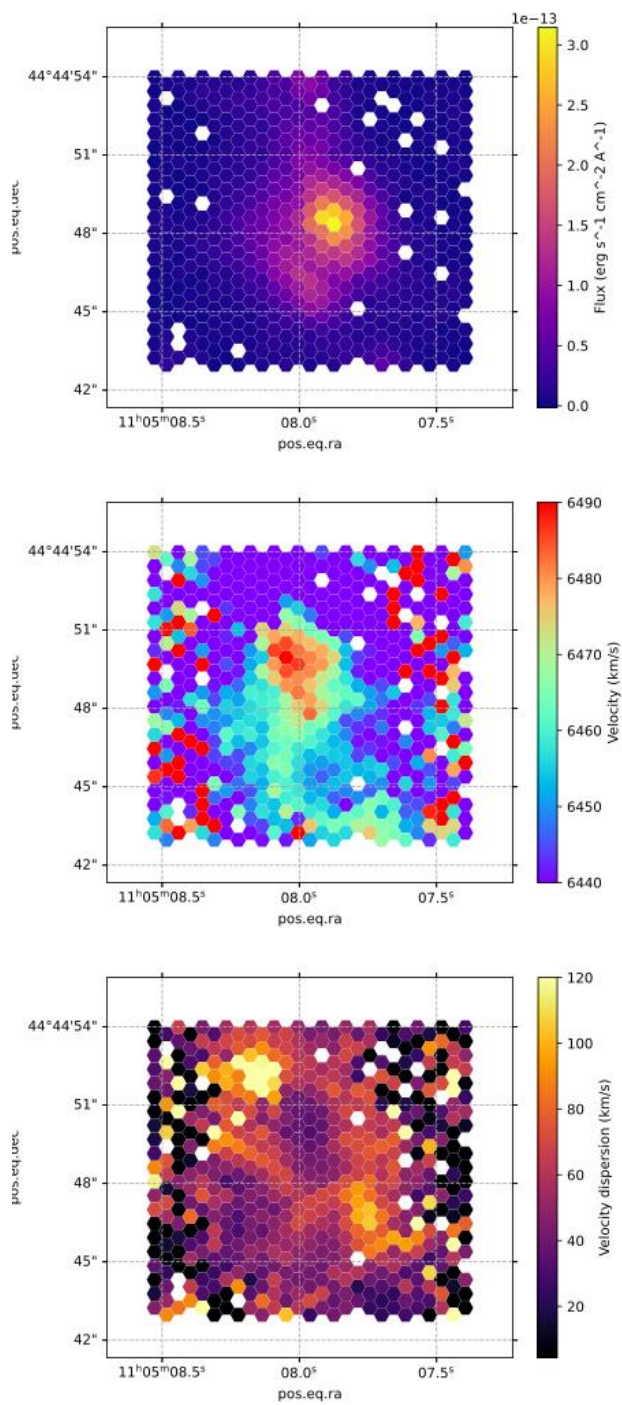
B system



[OII]3729

A system

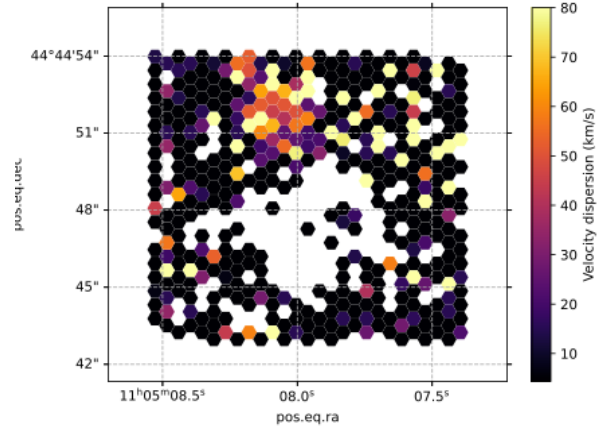
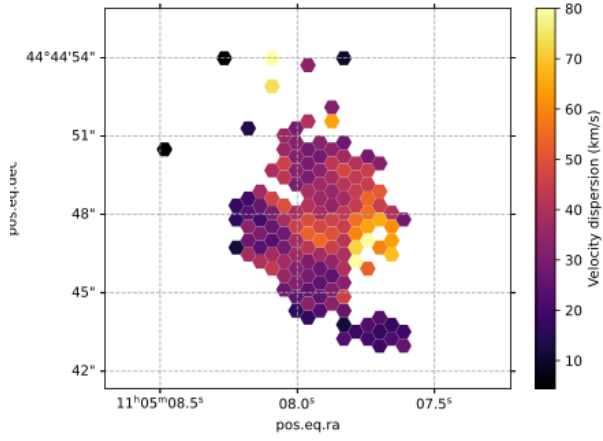
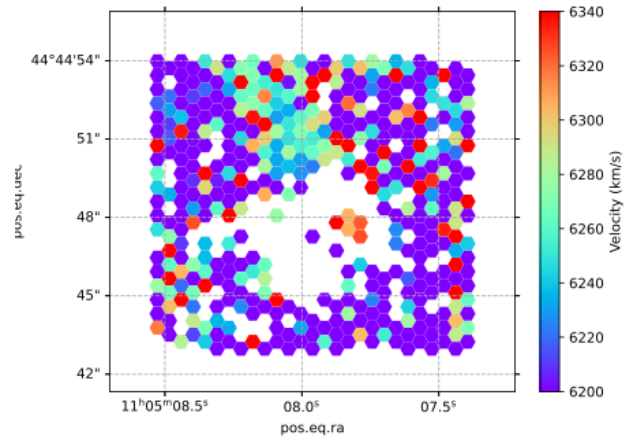
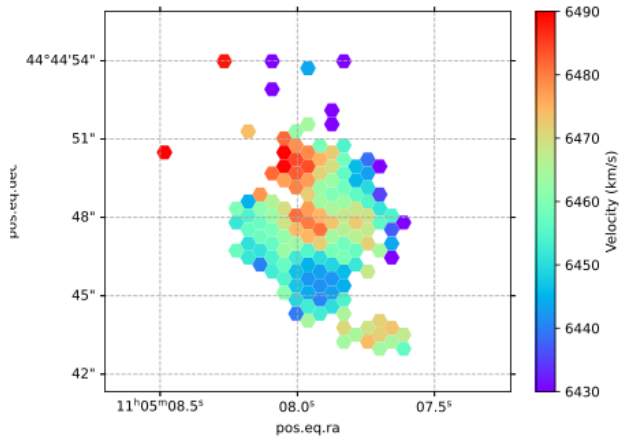
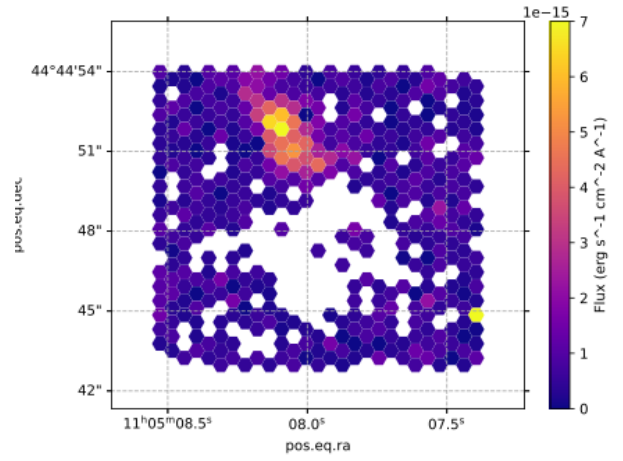
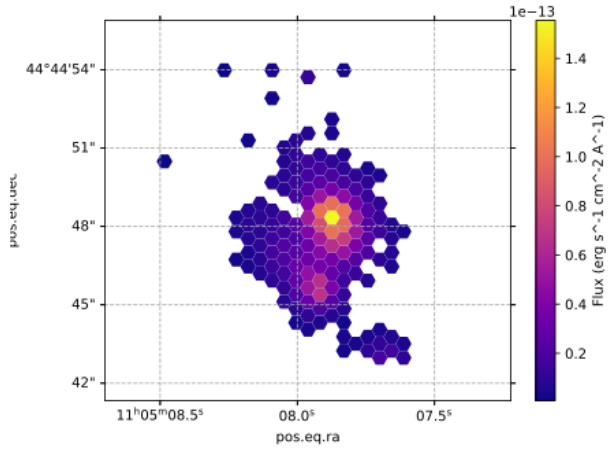
B system



[NeIII]3869

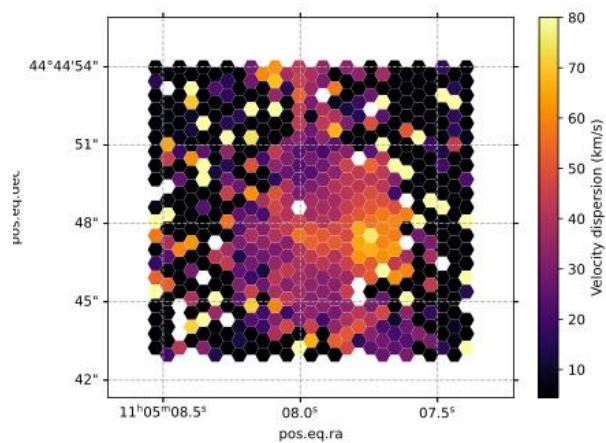
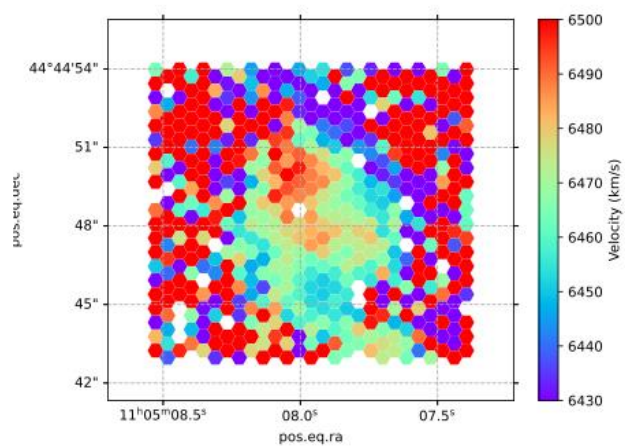
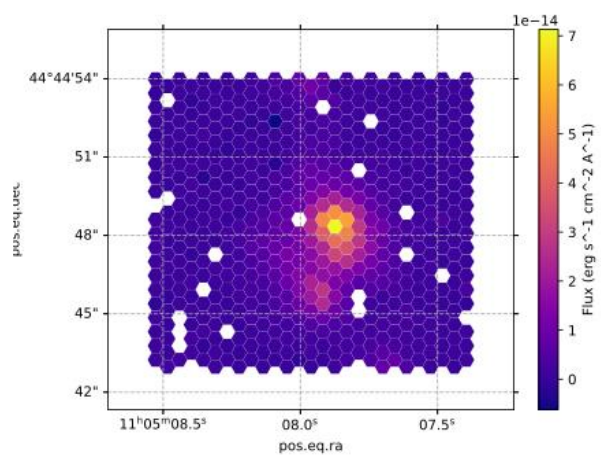
A system

B system

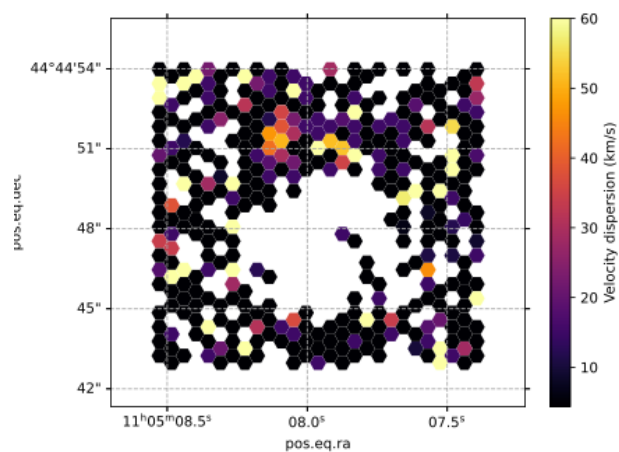
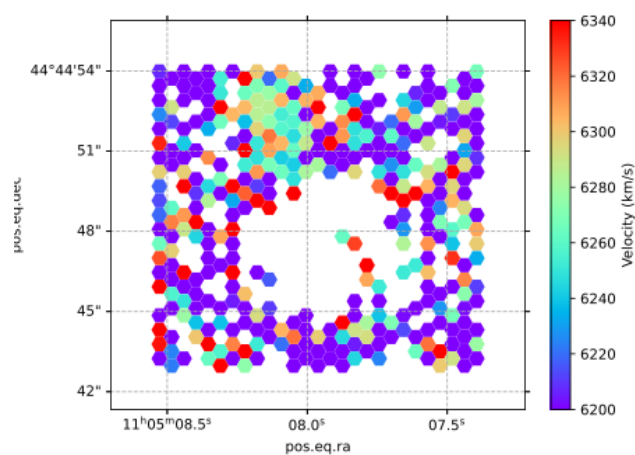
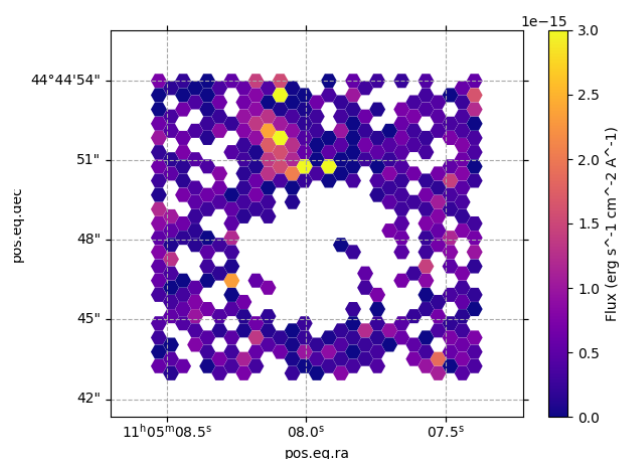


HeI3889

A system



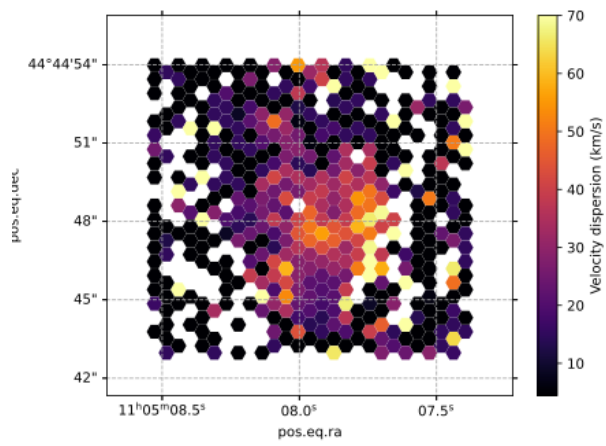
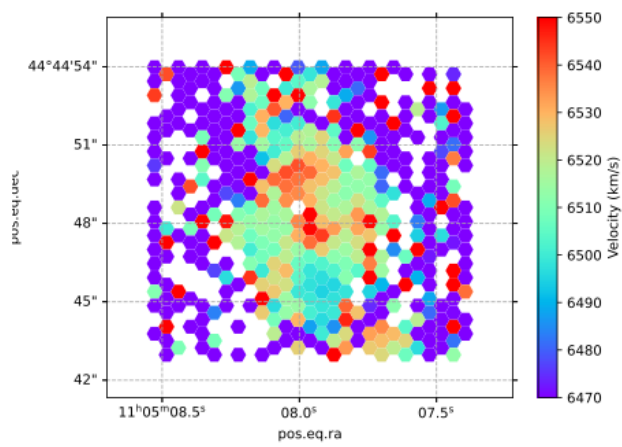
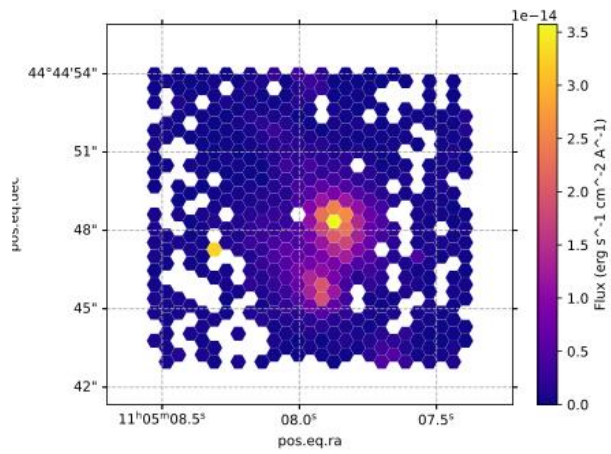
B system



[NeIII]3967

A system

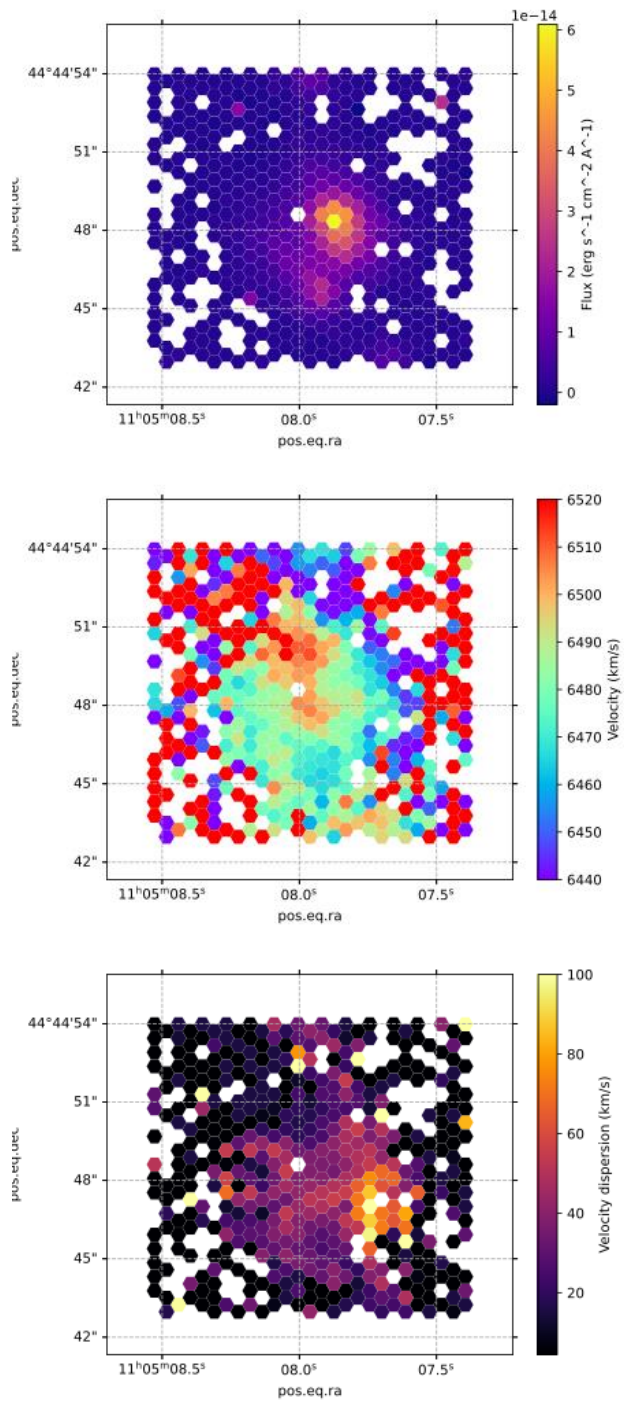
B system



H ϵ

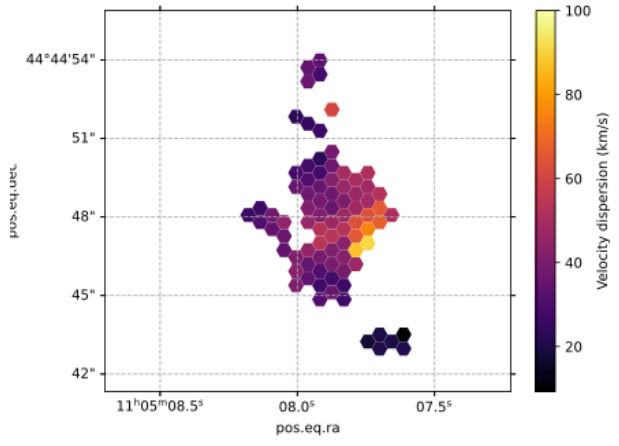
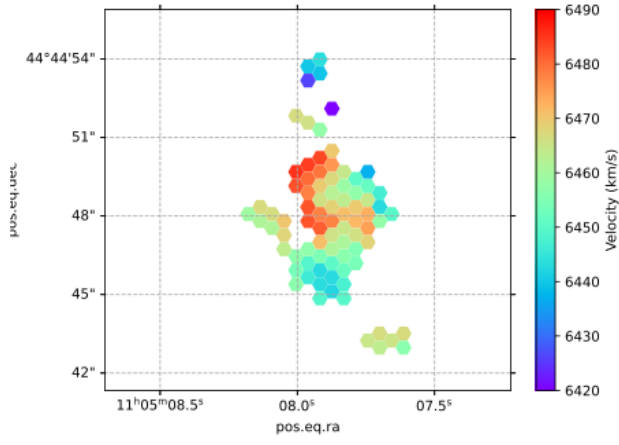
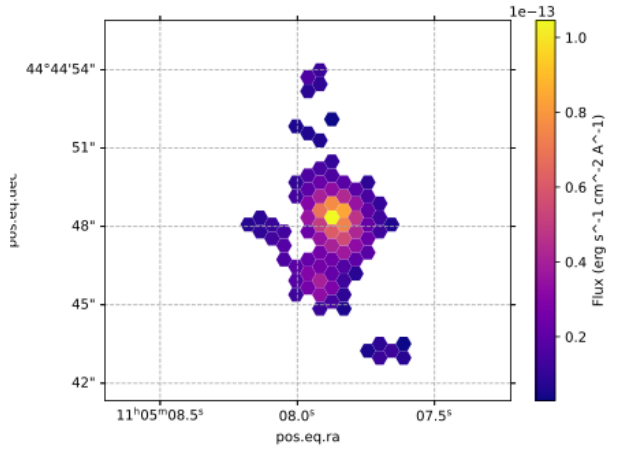
A system

B system

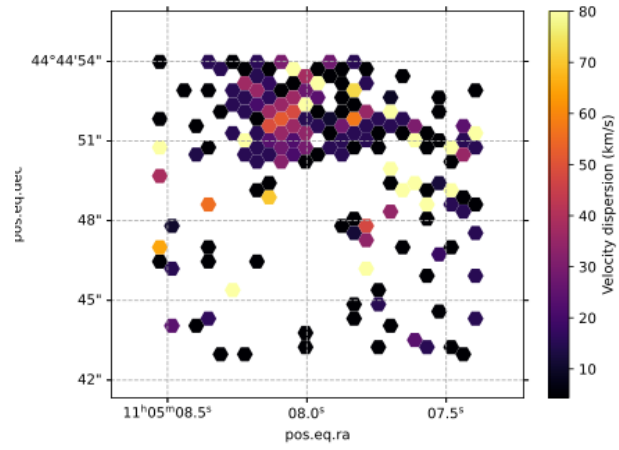
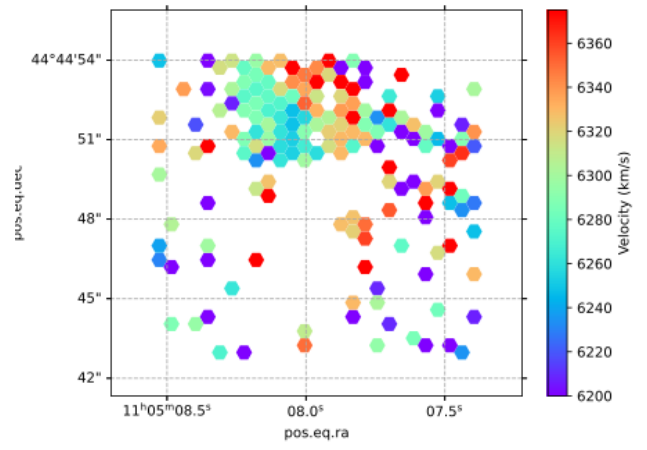
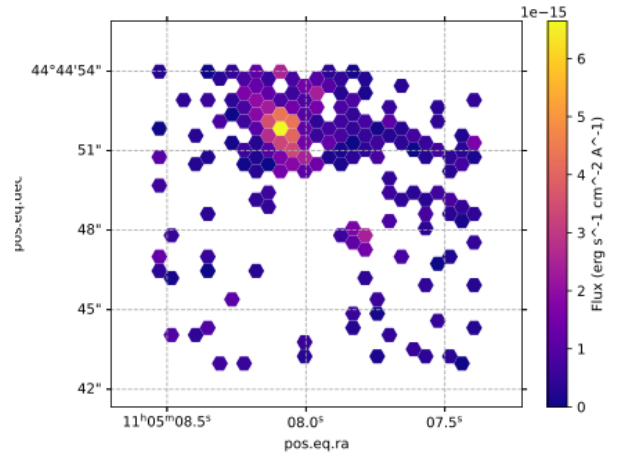


H δ

A system

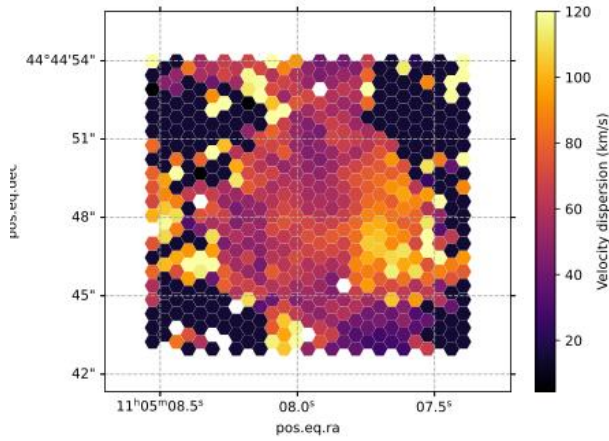
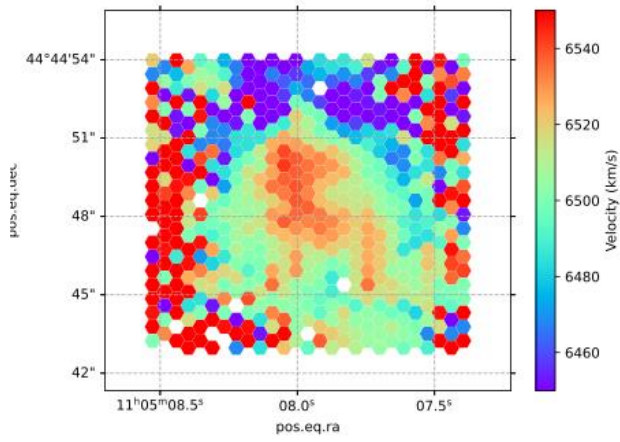
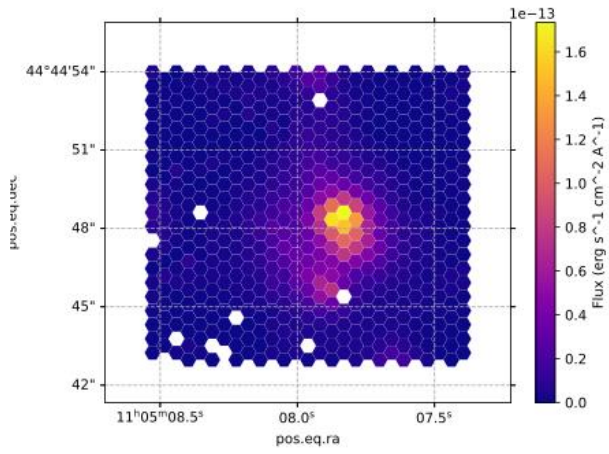


B system

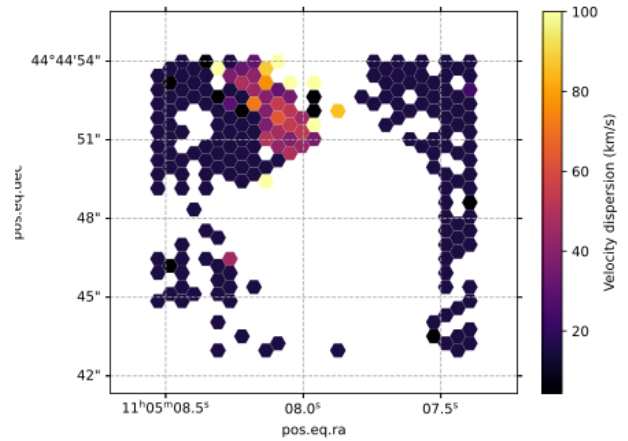
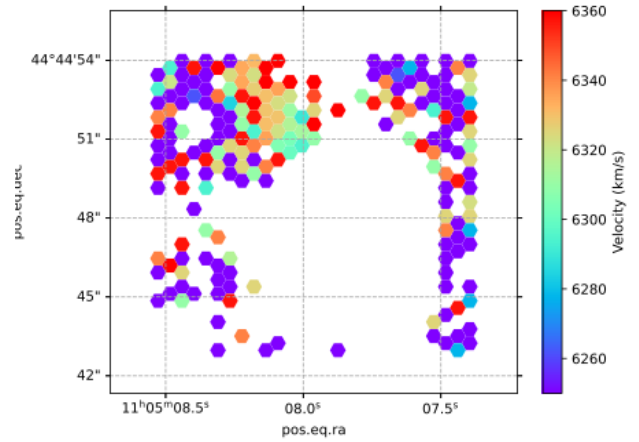
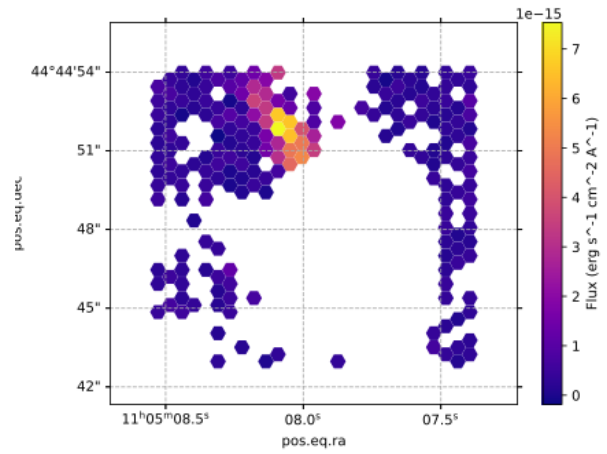


H γ

A system



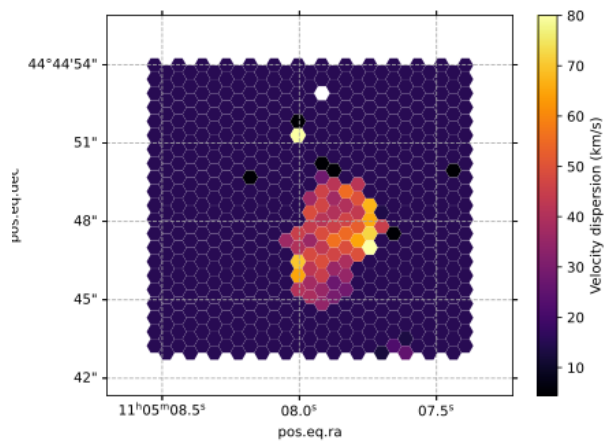
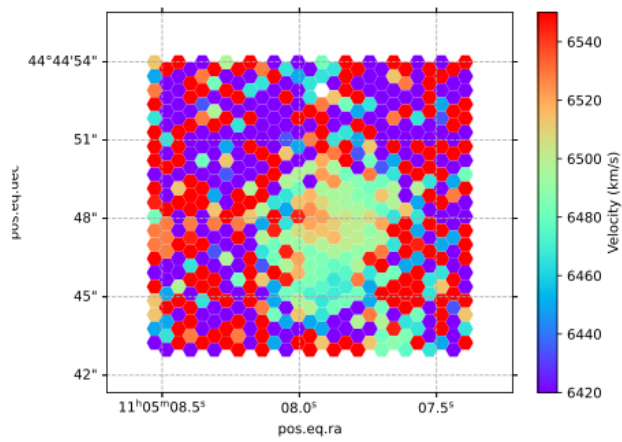
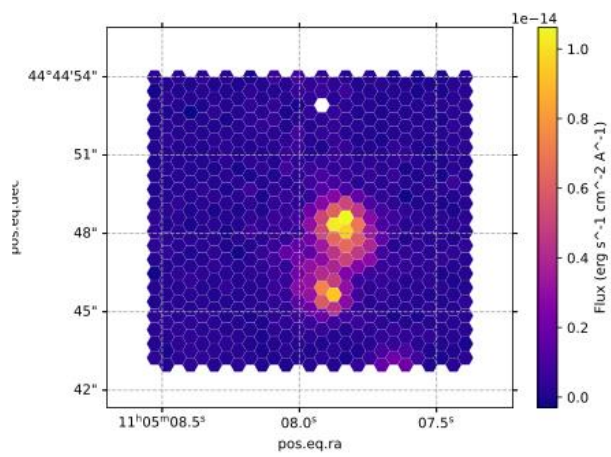
B system



[OIII]4363

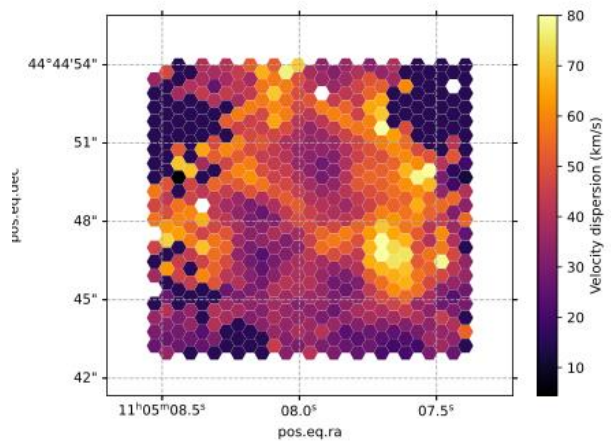
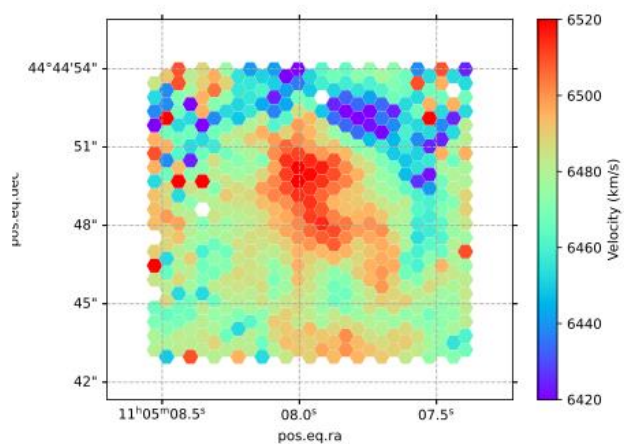
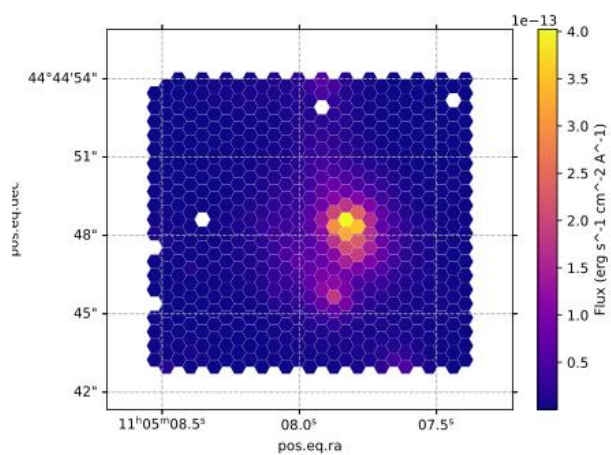
A system

B system

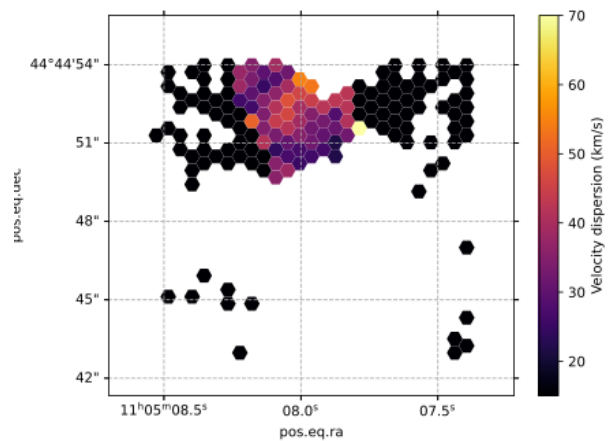
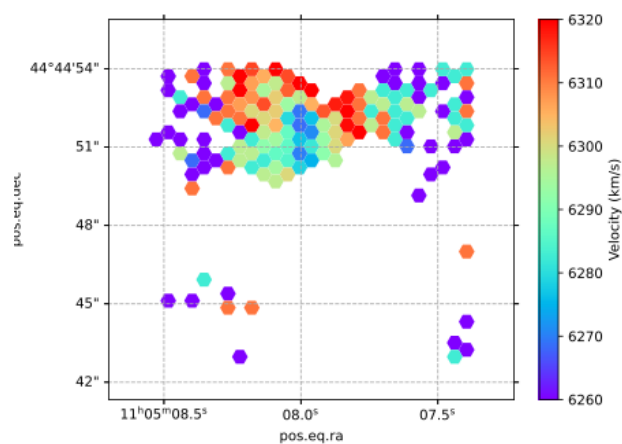
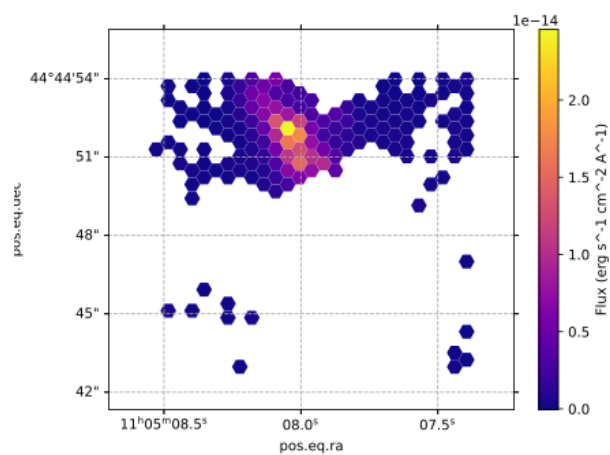


H β

A system



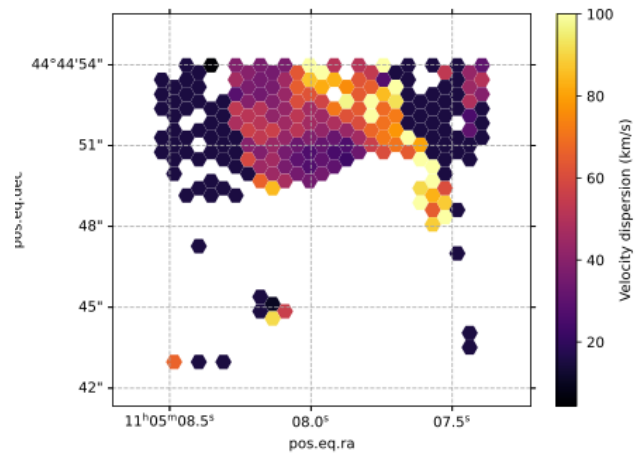
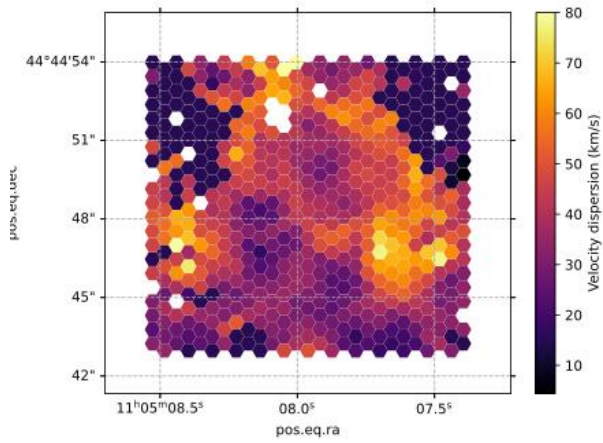
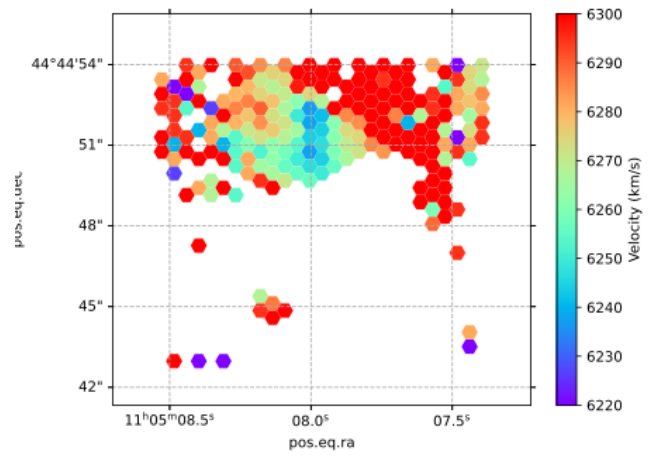
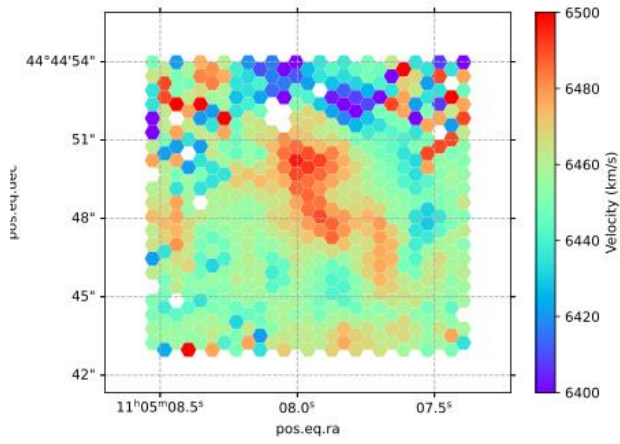
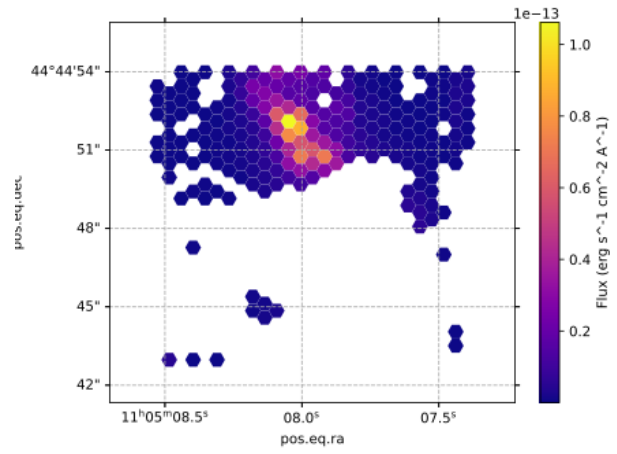
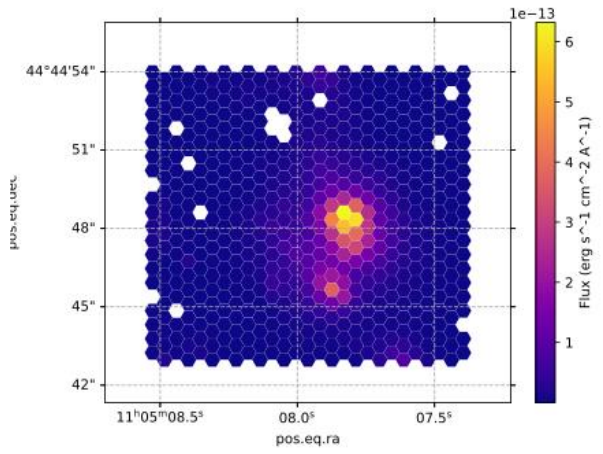
B system



[OIII]4959

A system

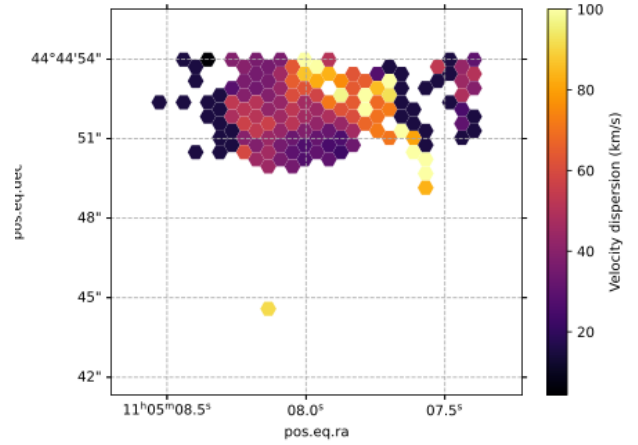
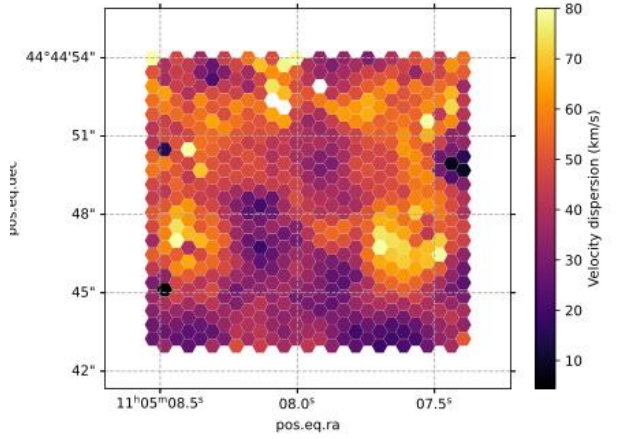
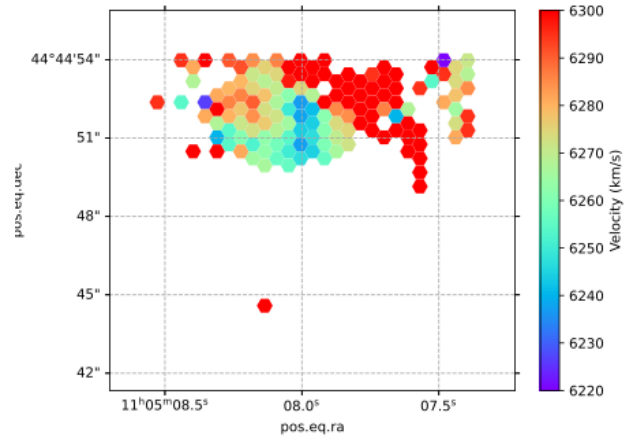
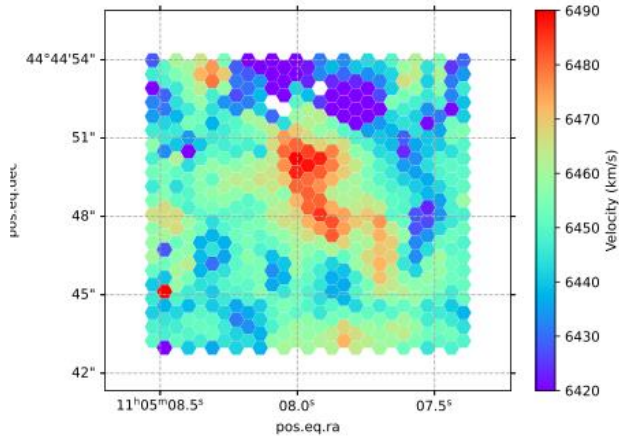
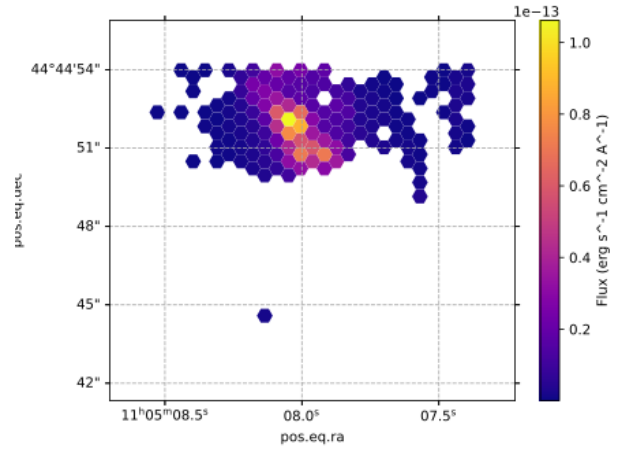
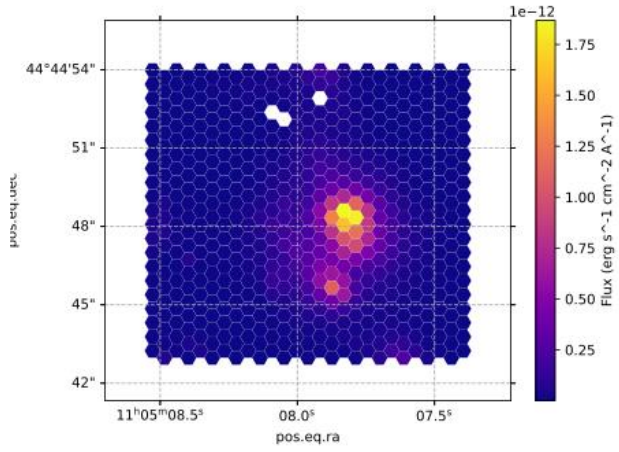
B system



[OIII]5007

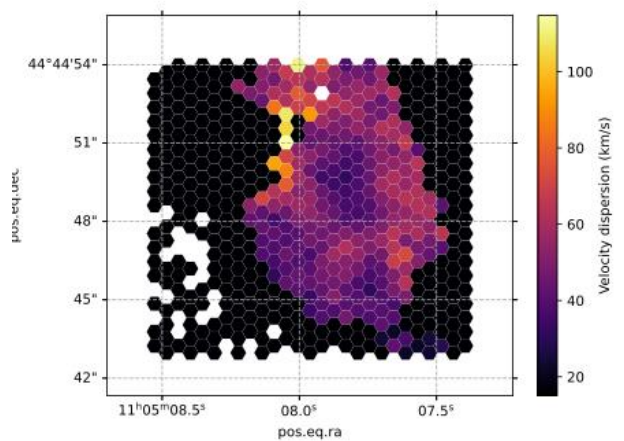
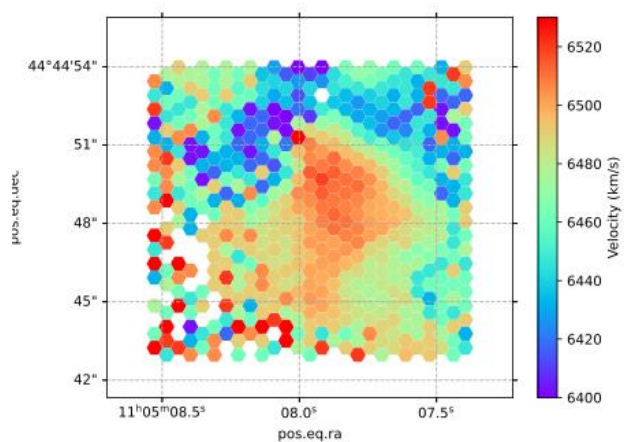
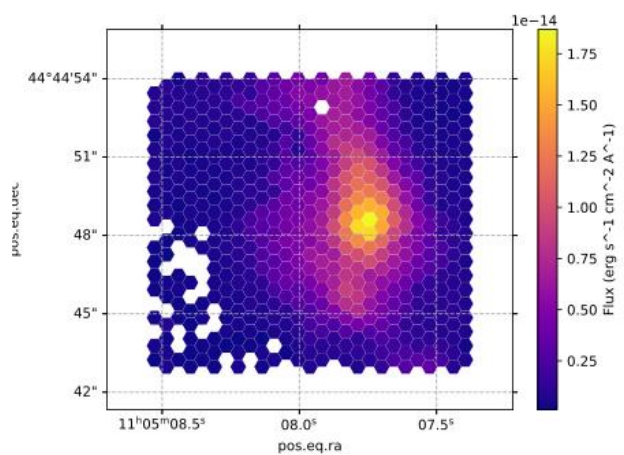
A system

B system

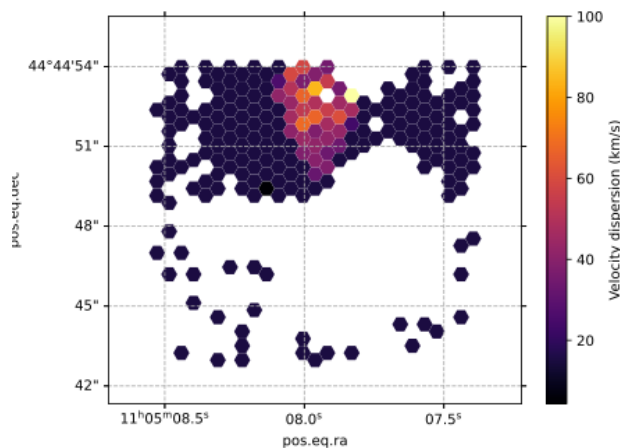
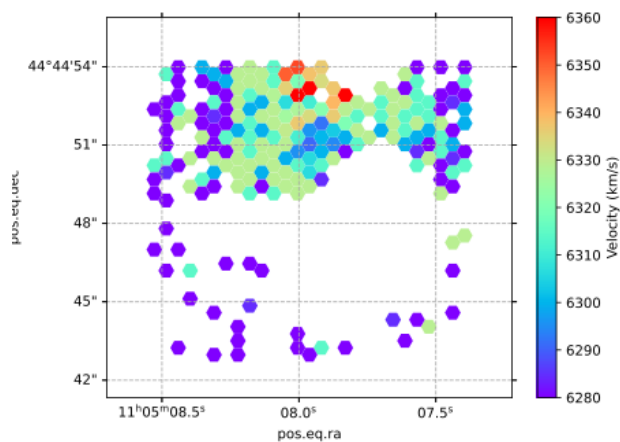
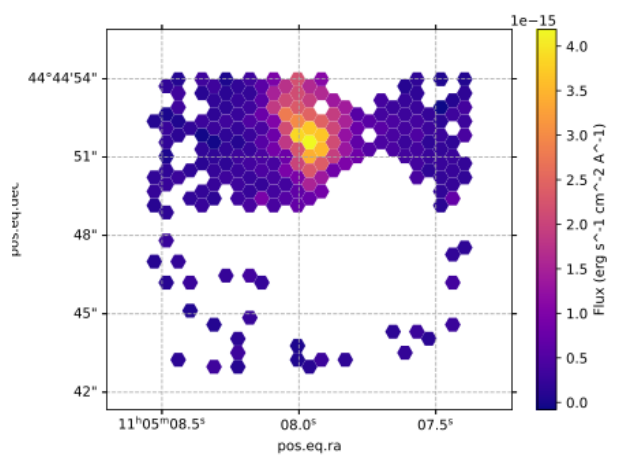


[OI]6300

A system

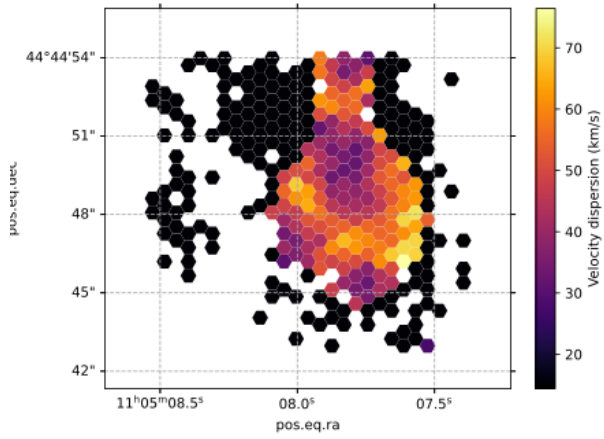
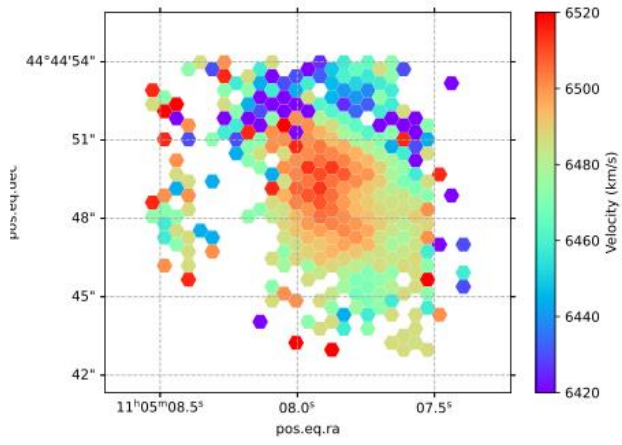
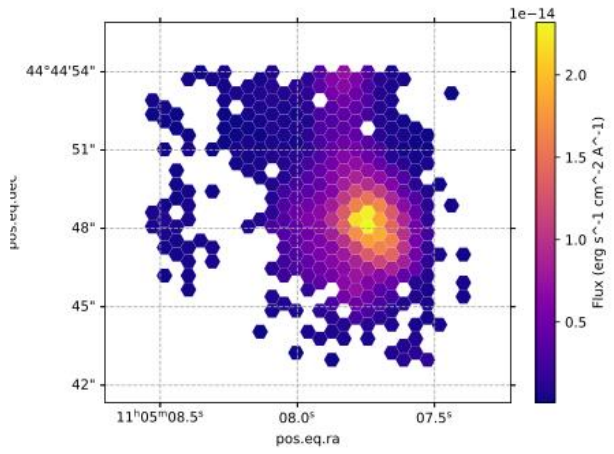


B system

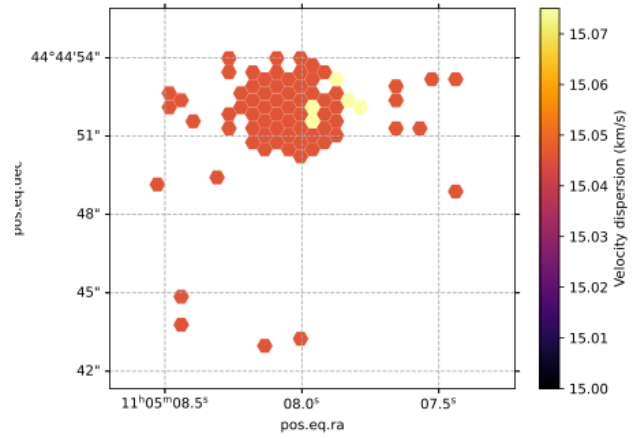
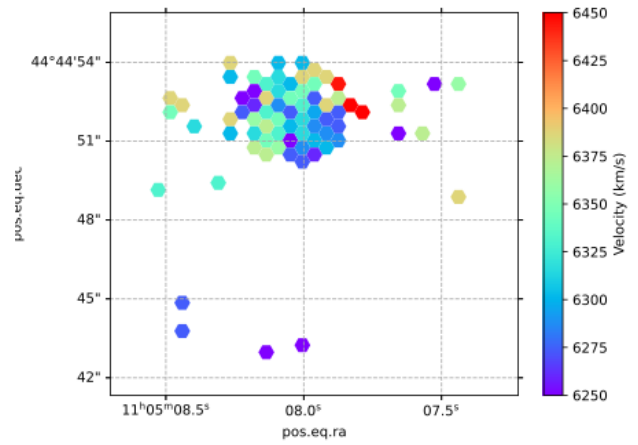
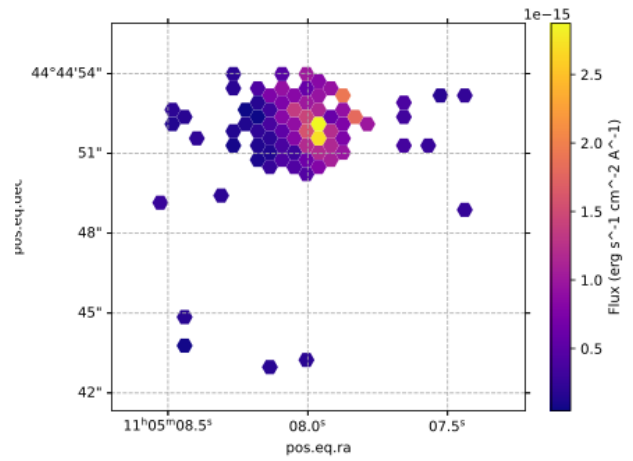


[NII]6548

A system

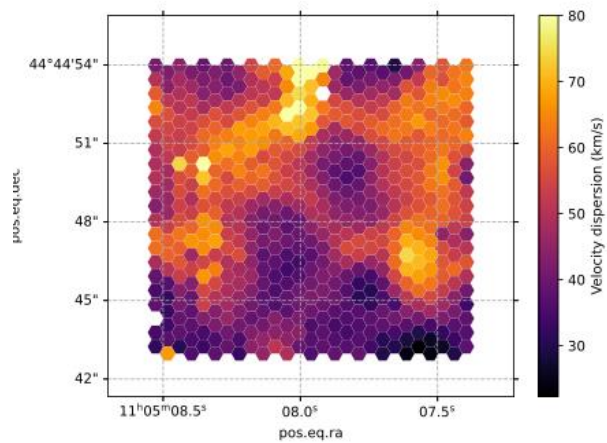
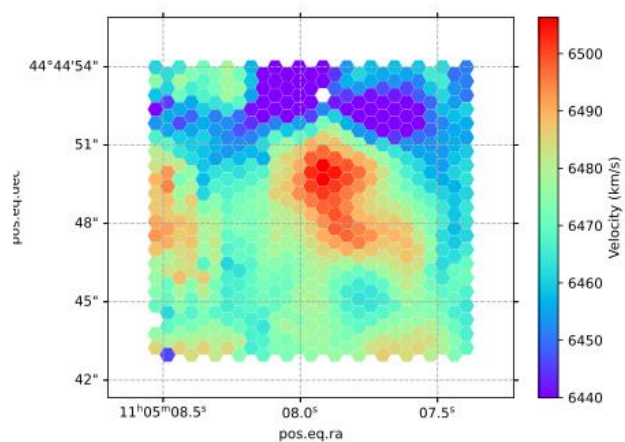
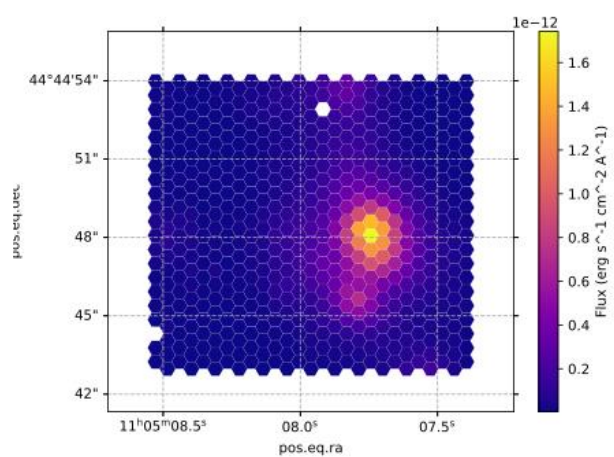


B system

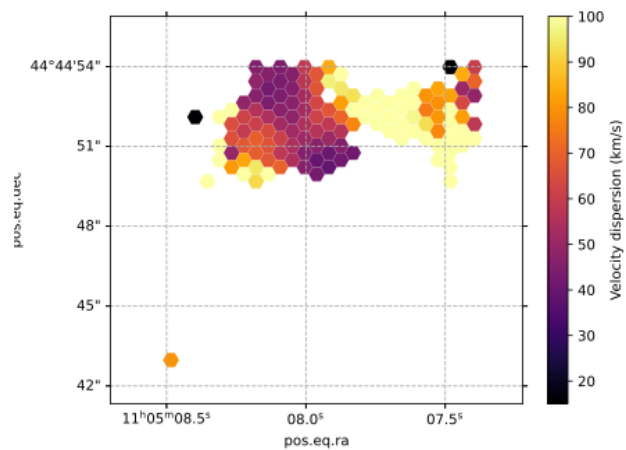
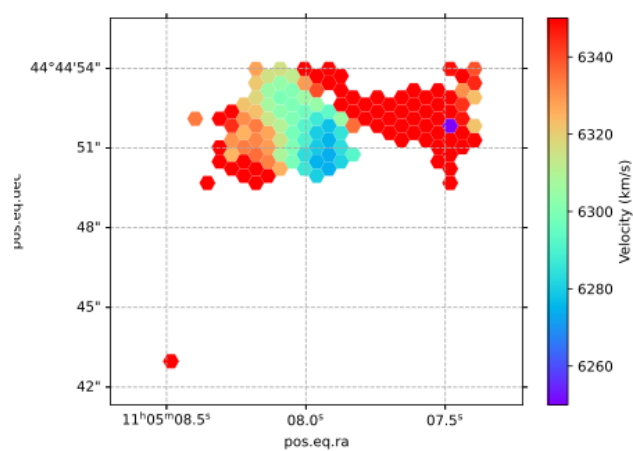
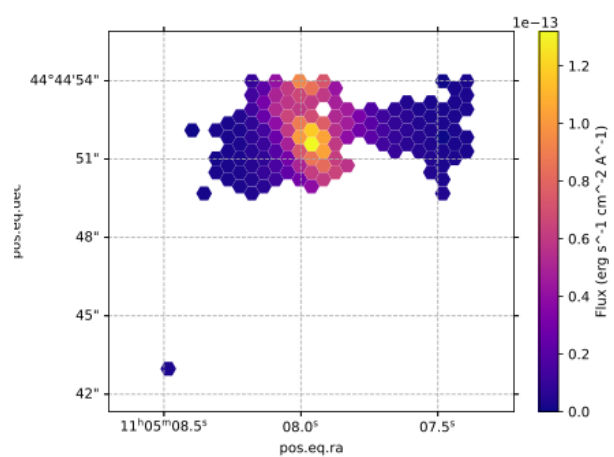


H α

A system



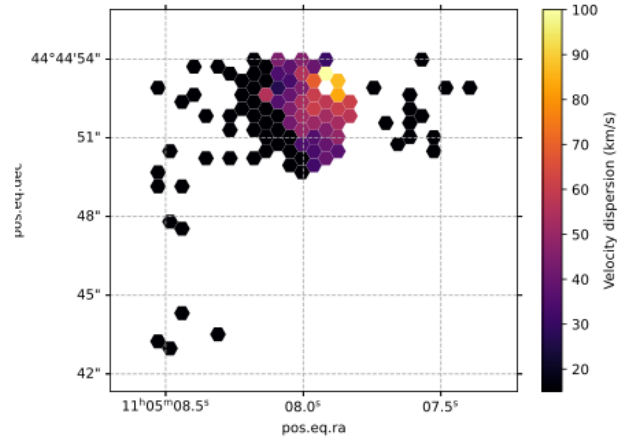
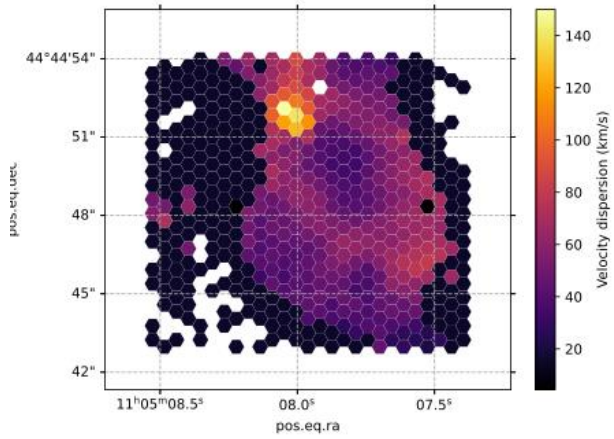
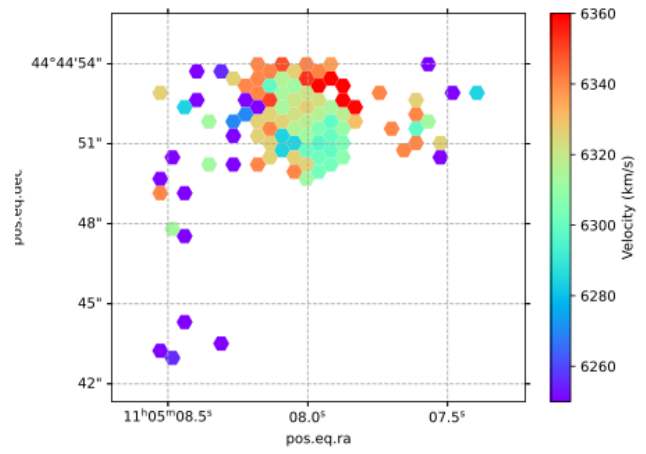
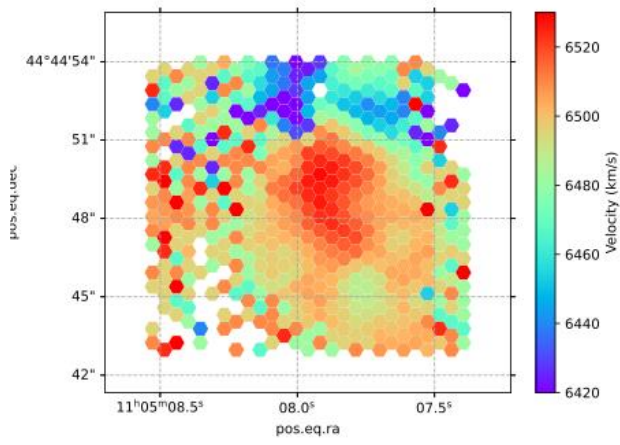
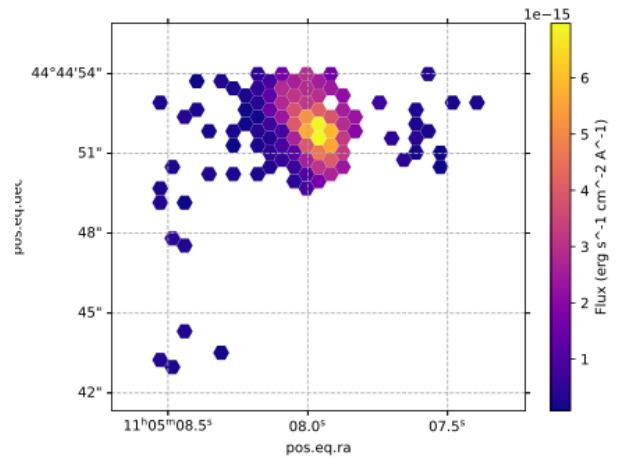
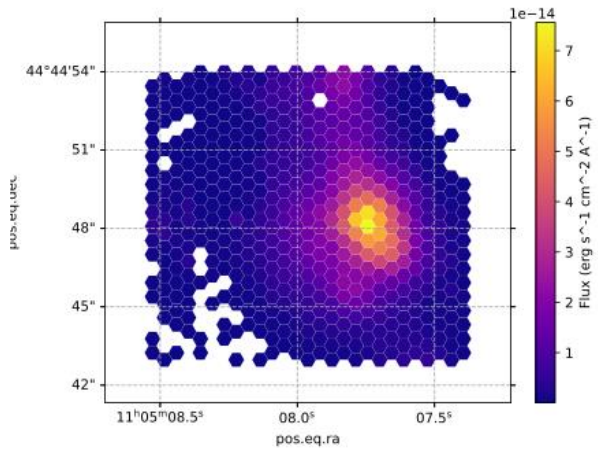
B system



[NII]6583

A system

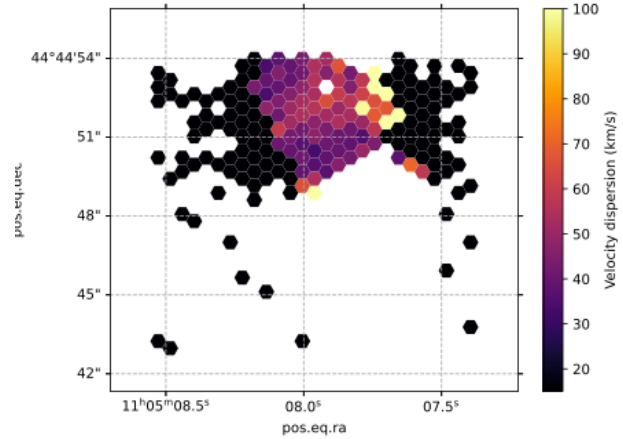
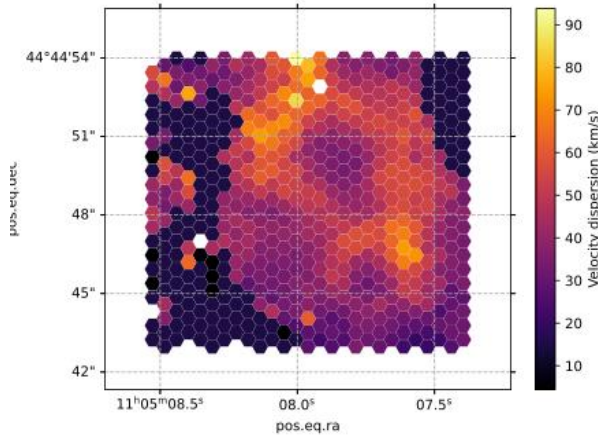
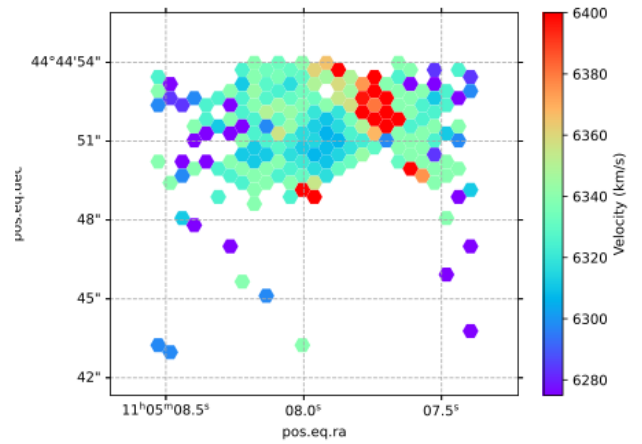
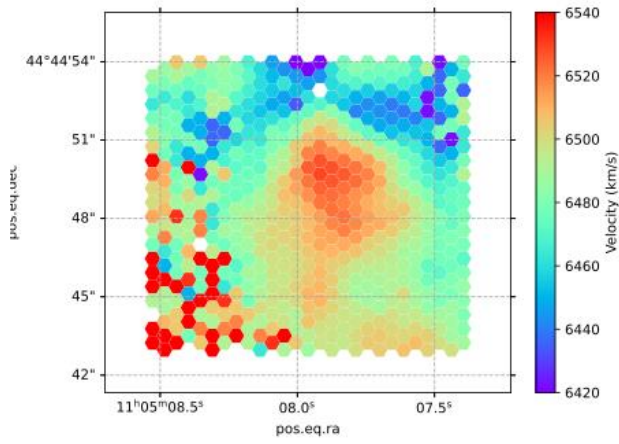
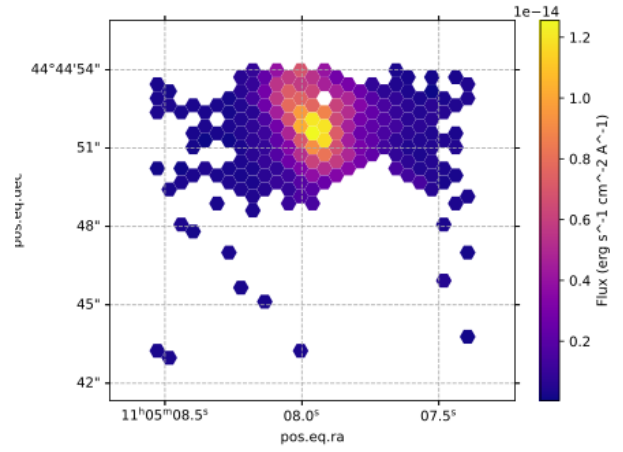
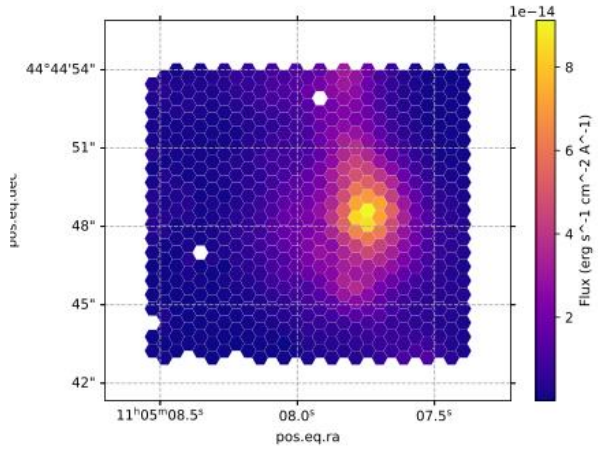
B system



[SII]6716

A system

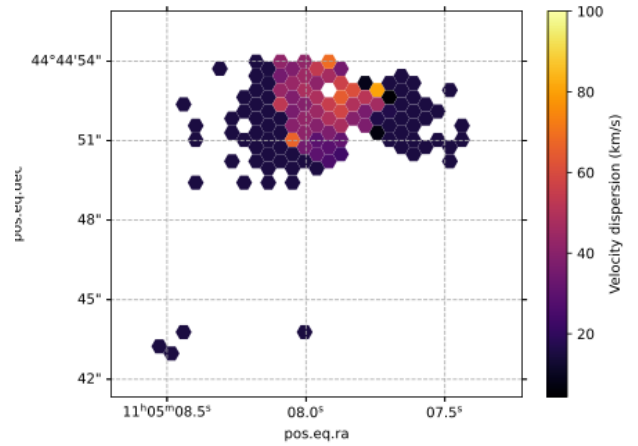
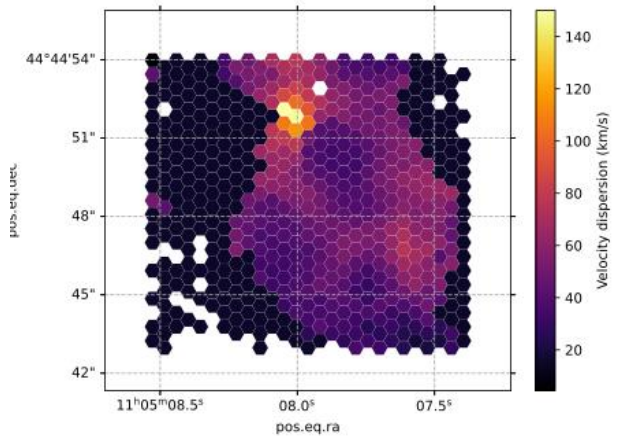
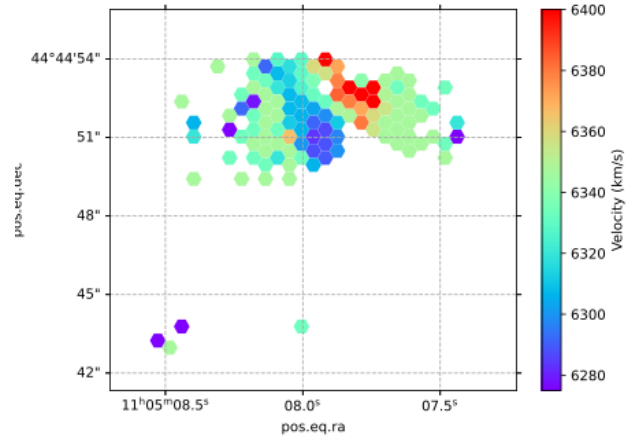
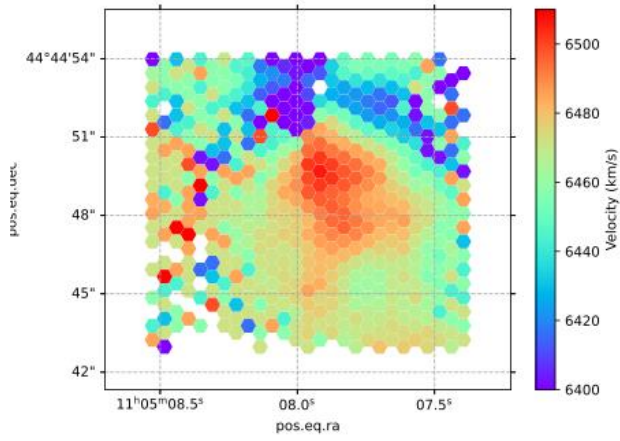
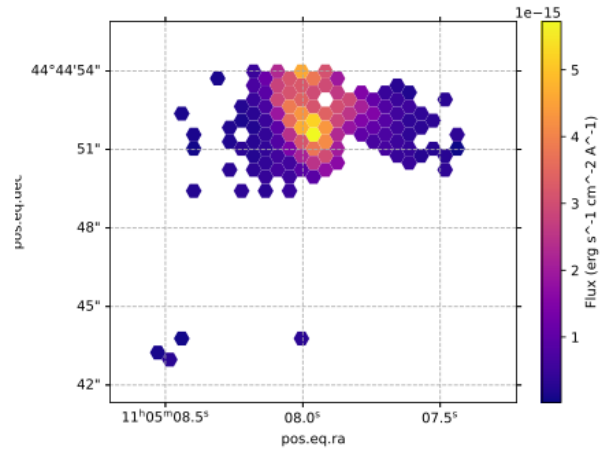
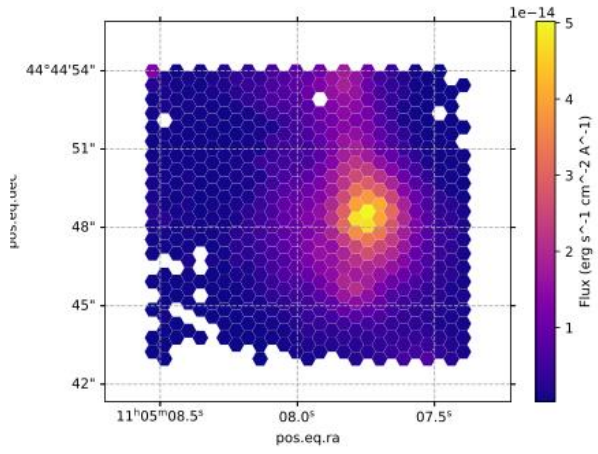
B system



[SII]6731

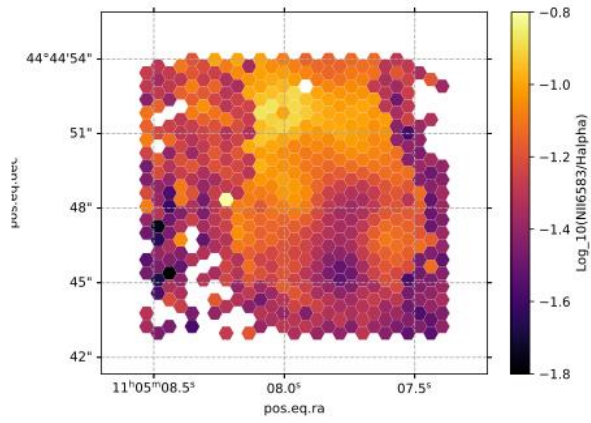
A system

B system

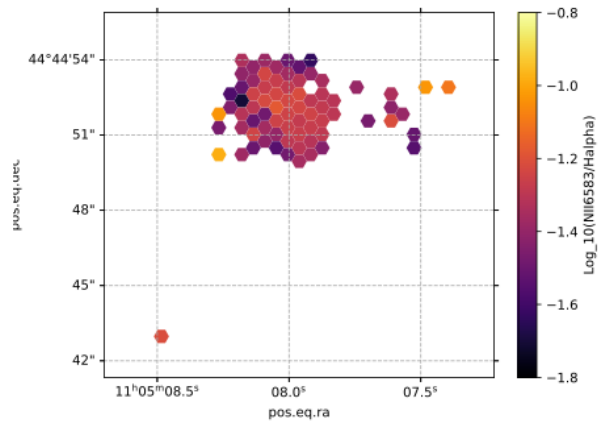


$$\log_{10}(\text{NII}6583/\text{H}\alpha)$$

A system

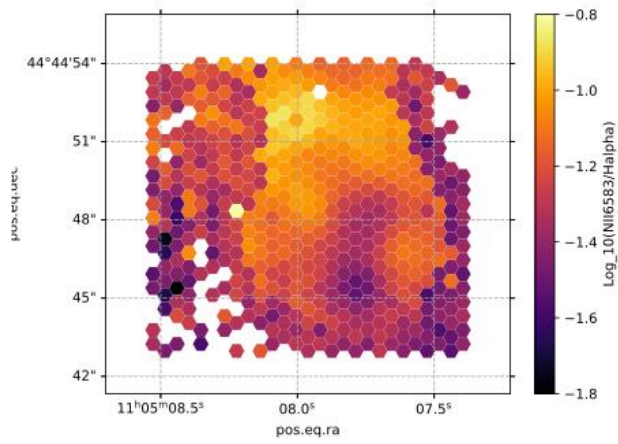


B system

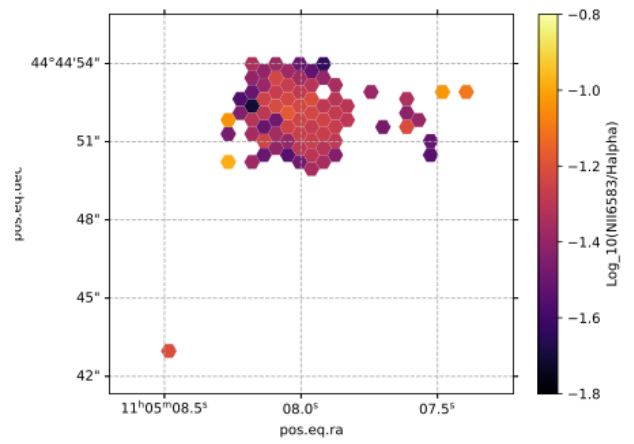


$$\log_{10}(\text{NII}6583/\text{NII}6548)$$

A system

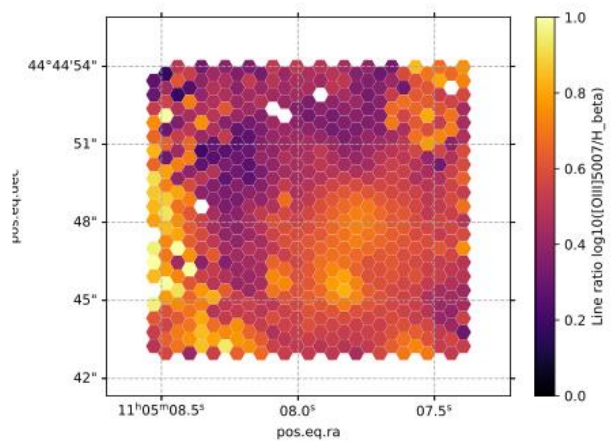


B system

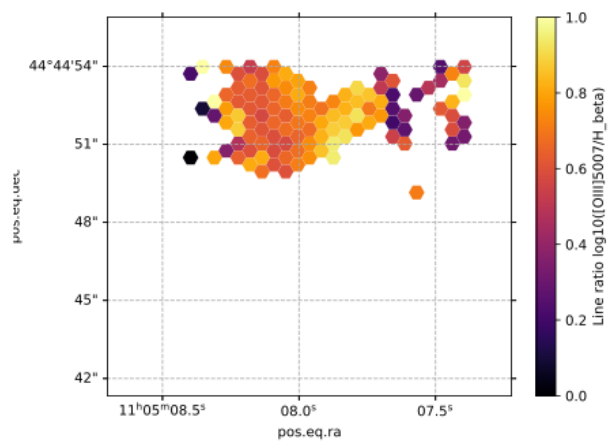


$$\log_{10}([\text{OIII}]\lambda 5007/\text{H}\beta)$$

A system

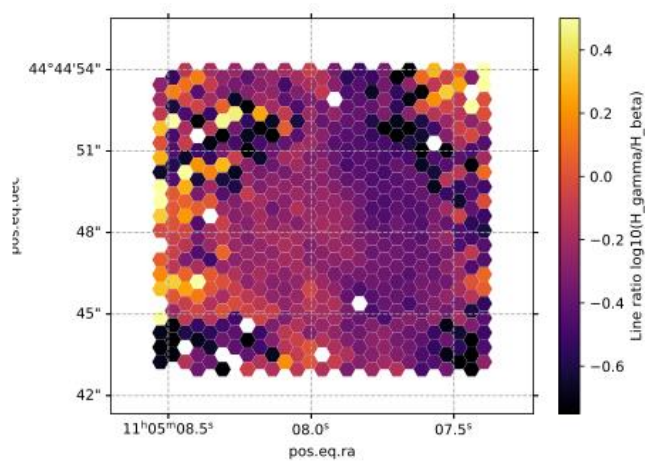


B system



$$\log_{10}(\text{H}\gamma/\text{H}\beta)$$

A system



B system

